

**The properties of plasma flows in fine ray structures
of the streamer belt as deduced from LASCO data**

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It is shown that within distances from the Sun's surface that are smaller than the height of the streamer helmet, each of two neighboring rays of the belt, as they approach the solar surface, bends around the helmet on either side thereof. Also, a minimum angular diameter of the rays of $d \approx 2^\circ - 3^\circ$ remains virtually constant within $R = (1.2 - 6.0)R_\odot$. A density inhomogeneities ("blob") can be produced above the helmet top to within $R \approx 6R_\odot$ at least. In this case the initial velocity of the "blob" increases with the solar distance, at which it is generated, and is something like the velocity of the bulk solar wind, by which the "blob" is carried away.

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